



SLR and the Next Generation Global Geodetic Networks

Erricos C. Pavlis & Magdalena Kuźmicz-Cieślak

Joint Center for Earth Systems Technology University of Maryland, Baltimore County and NASA Goddard, 698

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16th International Workshop on Laser Ranging

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- Space techniques are indispensable for the development of the terrestrial reference frame and for geodetic metrology
- The current state-of-the-art does not meet science requirements due to poor area
 coverage and aging equipment
- To meet the stringent future requirements (e.g. GGOS), we need to design a new network and deploy modern hardware systems









Outline



- SLR network
 - Present status
 - Future developments
- SLR contribution to ITRF
 - Accuracy assessment
 - Next generation TRF goals
- Simulations for network optimization
 - SLR & VLBI case studies
 - 8, 16, 24 and 32-site network results
- The next phase
 - Taking advantage of large & fast computer clusters (NASA's Columbia grid) for targeted test cases









Multiple techniques to solve the puzzle

- High precision geodesy is very challenging
 - 0.1 mm/yr stability required for sea level monitoring
- Fundamentally different observations with unique capabilities
- Together provide redundancy, cross validation and increased accuracy for TRF
- Strength from improvement of techniques and integration of techniques
- Fundamental prerequisite: Well-distributed, co-located networks with accurate ties

Technique Signal Source Obs. Type	VLBI Microwave Quasars Time difference	SLR Optical Satellite Two-way range	GPS Microwave Satellites Carrier phase
Celestial Frame UT1	<u>Yes</u>	No	No
Scale	<u>Yes</u>	<u>Yes</u>	Yes
Geocenter	No	<u>Yes</u>	Yes
Geographic Density	No	No	<u>Yes</u>
Real-time	No	No	Yes
Decadal Stability	<u>Yes</u>	<u>Yes</u>	Yes











- Single photon operational regime
- Narrow laser divergence
- Multi-kiloHertz operation (with multiple fires in flight)
- Autonomous, independent operations
- Improved epoch timing
- More stable / better defined pointing and ranging calibrations
- Eye-safe operation, LEO to GNSS
- Predictions and collected data submission via WWW (near real-time)
- Some new applications :
 - kHz scanning of satellite surface (allows for determination of spin-axis and rate);
 - Atmospheric seeing measurements along laser beam;
 - kHz Time Transfer (test using AJISAI and Graz system);
 - kHz LIDAR (under implementation now in Graz);
 - Detection of atmospheric layers, clouds, aircraft vapor trails;









LARES _{A/m} = 0.36 x LAGEOS



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Uncertainties due to Limited Knowledge or Modeling *NOW* 5-10 mm 1-5 mm 1-5 mm **Improvements:** Improved s/c CoM offsets New refraction modeling with gradients **Atmospheric Loading & Gravitational Potential** Better ground survey and eccentricity monitoring 10-30 mm Copyright 2006 © Teddy Pavlis









Design of the Future Network

- SLR and VLBI optimal combination (first step):
 - ✓ Simulate SLR and VLBI data for 2004 from four networks of 8, 16, 24 and 32 sites
 - ✓ Assume system performance of NGSLR and VLBI2010
 - Simulation of a 1-year period with SLR and VLBI data (eventually to be extended to ~ 6 years)

– Inclusion of GNSS, DORIS, etc. later, in a future step







Simulation Goal

Which network will deliver consistently and reliably:

<1 mm epoch position and < 0.1 mm/y secular change</p>

Inclusion factor clied (neclarang 1 and in 1 million years)









- Primarily a test to verify the simulation process end-to-end •
- Four networks with 8, 16, 24 and 32 sites •
- Only site positions and EOP estimated from one year of data ۲
- Scaled error covariance projected on the 7 TRF parameters •
- Assuming that errors across years are uncorrelated, we project ٠ the one year results to estimate the number of years to reach our accuracy goals







Network variants (8 \Rightarrow 32)







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One-year Simulation Results







SLR+VLBI_sim8_080322









One-year Simulation Results

Origin & Scale **SLR ONLY** 50 ► 24 vs. 8 - 32 vs. 8 -24 vs. 16 40 🔫 -32 vs. 16 Relative Improvement [%] - 2 · 32 vs. 24 30 20 10 0 Ту Tx Τz Scale **Network Similarity Parameters** SLR+VLBI_sim8_080322scIMM



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One-year Simulation Results



- The simulation validates the real world experience with 8 sites
- The biggest improvement is seen when going from 8 to 16 sites
- The largest impact of an 8 site addition in the origin is seen when going from 16 to 24 sites (~22%), and the least, from 24 to 32 (~8%)
- Results for a 13 year time span (corresponds to ITRF2005) show a 4- or 5-fold improvement compared to what we estimate for ITRF2005
- A projection for a 16 year time span (*ITRF2008*?) shows that a 32 site network approaches the GGOS goal of accuracy in the origin and scale







Some Simulation Issues:



- We currently work with two techniques only (SLR & VLBI)
- Optimal network size with constrained system performance and background model quality
- Assuming perfect site-ties
- Criterion is "TRF" quality: origin, scale and orientation
- Need to consider temporal variations of the TRF parameters
- Solutions to be repeated with the addition of local tie errors with varied weighting schemes
- We will use the 16 site network to investigate the effect of choosing alternate sites on the results (varying the uniformity of the network)







Summary



- Origin and scale marginally controlled by a 24 site network; when extended to 32 sites, it approaches GGOS goals (1 mm)
- Orientation seems to be less dependent on the size of the network
- The effect of additional techniques on the quality of the TRF remains to be assessed
- Need to develop scenarios of "degradation" and "improvement" of nominal design parameters







Future Work



- We may have to consider *improvement of our* models, analysis techniques and our space segment (e.g. SLR targets) to improve TRF accuracy while keeping a reasonable network size to reach our goal
- Our simulation process now runs on a faster CPU to allow a quicker turn-around of future cases (Columbia grid cluster)
- As we improve our turn-around time we plan to investigate scenarios with additional parameters varied (more satellites, different orbits, systematic errors, operational modes, etc.)







<1 mm epoch position and < 0.1 mm/y secular variations













Back-up slides

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Maximal Overlapping SLR-VLBI Network [32]

Next Generation NASA Networks ~70 sites





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Why 1 mm / 0.1 mm/y?

For every 1 mm/y Z-trend in



ITRF2005: 3.3 +/- 0.07 mm/yr



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radial = -0.17 mm/vr

2005

2005.5

2004.5



Subset Solutions for an SLR TRF







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TRF Subset Solutions Statistics [mm]

Case	ΔΧ σ _{ΔΧ}	ΔΥ σ _{ΔΥ}	ΔΖ σ _{ΔΖ}	3D Δ σ _{3D Δ}
3 Odd	-8.37 ±10.91	19.25 ±10.78	-4.20 ±10.32	21 ± 17
4 Even	-12.62 ± 8.93	5.15 ± 8.82	-12.50 ± 8.44	18 ± 16
1 1/2	-41.20 ±35.82	6.26 ±35.38	-10.10 ±33.86	43 ± 61
2	1.74 ± 6.76	8.06 ± 6.68	7.28 ± 6.39	11 ± 11
15 1/4	-60.49 ± 23.68	57.43 ±23.39	7.48 ±22.39	84 ± 40
16	18.65 ± 31.40	-57.81 ±30.88	-6.19 ±29.50	61 ± 53
17	-0.27 ± 18.01	-4.74 ±17.79	15.72 ±17.03	16 ± 31
18	2.07 ± 12.29	7.16 ±12.18	1.73 ±11.60	8 ± 21



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